

# Single-particle potential in a chiral approach to nuclear matter including short-range NN-terms

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**Abstract.** We extend a recent chiral approach to nuclear matter of Lutz *et al.* (Phys. Lett. B **474**, 7 (2000)) by calculating the underlying (complex-valued) single-particle potential  $U(p, k_f) + iW(p, k_f)$ . The potential for a nucleon at the bottom of the Fermi sea,  $U(0, k_{f0}) = -20.0$  MeV, comes out as much too weakly attractive in this approach. Even more seriously, the total single-particle energy does not rise monotonically with the nucleon momentum  $p$ , implying a negative effective nucleon mass at the Fermi surface. Also, the imaginary single-particle potential,  $W(0, k_{f0}) = 51.1$  MeV, is too large. More realistic single-particle properties together with a good nuclear-matter equation of state can be obtained if the short-range contributions of non-pionic origin are treated in mean-field approximation (*i.e.* if they are not further iterated with  $1\pi$ -exchange). We also consider the equation of state of pure neutron matter  $\bar{E}_n(k_n)$  and the asymmetry energy  $A(k_f)$  in that approach. The downward bending of these quantities above nuclear-matter saturation density seems to be a generic feature of perturbative chiral pion-nucleon dynamics.

**PACS.** 12.38.Bx Perturbative calculations – 21.65.+f Nuclear matter

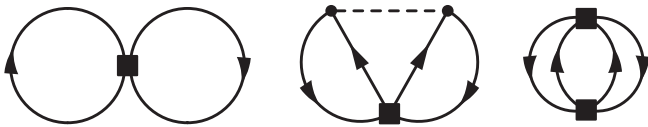
## 1 Introduction

The present status of the nuclear-matter problem is that a quantitatively successful description can be achieved, using advanced many-body techniques [1], in a non-relativistic framework when invoking an adjustable three-body force. Alternative relativistic mean-field approaches, including non-linear terms with adjustable parameters or explicitly density-dependent point couplings, are also widely used for the calculation of nuclear-matter properties and finite nuclei [2].

In recent years a novel approach to the nuclear-matter problem based on effective field theory (in particular chiral perturbation theory) has emerged [3–5]. The key element there is a separation of long- and short-distance dynamics and an ordering scheme in powers of small momenta. At nuclear-matter saturation density the Fermi momentum  $k_{f0}$  and the pion mass  $m_\pi$  are comparable scales ( $k_{f0} \simeq 2m_\pi$ ), and therefore pions must be included as explicit degrees of freedom in the description of the nuclear many-body dynamics. The contributions to the energy per particle of isospin-symmetric nuclear matter  $\bar{E}(k_f)$  originating from chiral pion-nucleon dynamics have been calculated up to three-loop order in refs. [3,4]. Both calculations include the  $1\pi$ -exchange Fock diagram and the iter-

ated  $1\pi$ -exchange Hartree and Fock diagrams. In ref. [4] irreducible  $2\pi$ -exchange is also taken into account and a momentum cut-off  $\Lambda$  is used to regularize the few divergent parts associated with chiral  $2\pi$ -exchange. The resulting cut-off-dependent contribution to  $\bar{E}(k_f)$  is completely equivalent to that of a zero-range NN-contact interaction (see eq. (15) in ref. [4]). At that point the (earlier) work of Lutz *et al.* [3] follows a different strategy. Two zero-range NN-contact interactions (acting in  $^3S_1$  and  $^1S_0$  NN-states) proportional to the parameters  $g_0 + g_A^2/4$  and  $g_1 + g_A^2/4$  are introduced (see eq. (4) in ref. [3]). The components proportional to  $g_A^2/4$  cancel the zero-range contribution generated by the  $1\pi$ -exchange Fock diagram. The other components proportional to  $g_0$  and  $g_1$  are understood to subsume all non-perturbative short-range NN-dynamics relevant at densities around nuclear-matter saturation density  $\rho_0$ . In order to be consistent with this interpretation the NN-contact vertices proportional to  $g_{0,1}$  are allowed to occur only in first order. Furthermore, according to ref. [6] pions can be treated perturbatively (at least) in the  $^1S_0$  partial wave of NN-scattering if the zero-range pieces they generate are removed order by order. Therefore, the NN-contact vertex proportional to  $g_A^2/4$  occurs also in higher orders (see fig. 1 in ref. [3] which includes diagrams with “filled circle” and “open circle” vertices).

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**Fig. 1.** Additional in-medium diagrams generated by the NN-contact interactions introduced in ref. [3]. The two NN-contact interactions proportional to  $\gamma + 1$  and  $\gamma_n + 1$  are symbolized by the filled square vertex. The last diagram is to be understood such that quadratic terms (such as  $\gamma^2$ ,  $\gamma\gamma_n$  and  $\gamma_n^2$ ) are omitted.

Despite their differences in the treatment of the effective short-range NN-dynamics both approaches [3,4] are able to reproduce correctly the empirical nuclear-matter properties (saturation density  $\rho_0$ , binding energy per particle  $-\bar{E}(k_{f0})$  and compressibility  $K$ ) by adjusting only one parameter, either the coupling  $g_0 + g_1 \simeq 3.23$  or the cut-off  $\Lambda \simeq 0.65$  GeV. Note that in dimensional regularization all diagrams evaluated in ref. [3] are finite. In the chiral approach of the Munich group [4,5] the asymmetry energy  $A(k_f)$ , the energy per particle of pure neutron matter  $\bar{E}_n(k_n)$  as well as the (complex) single-particle potential  $U(p, k_f) + iW(p, k_f)$  below the Fermi surface ( $p \leq k_f$ ) have been calculated. Good results (in particular for the asymmetry energy,  $A(k_{f0}) = 33.8$  MeV, and the depth of the single-particle potential,  $U(0, k_{f0}) = -53.2$  MeV) have been obtained with the single cut-off scale  $\Lambda \simeq 0.65$  GeV adjusted to the binding energy per particle  $-\bar{E}(k_{f0}) = 15.3$  MeV. Moreover, when extended to finite temperatures [7] this approach reproduces the liquid-gas phase transition of isospin-symmetric nuclear matter, however, with a too high value of the critical temperature  $T_c = 25.5$  MeV.

It is the purpose of this work to investigate in the approach of Lutz *et al.* [3] the single-particle potential  $U(p, k_f) + iW(p, k_f)$  in isospin-symmetric nuclear matter, as well as the neutron matter equation of state  $\bar{E}_n(k_n)$  and the asymmetry energy  $A(k_f)$ . We conclude that the treatment of the effective short-range NN-interaction in that approach is insufficient in order to reproduce all empirical nuclear-matter properties.

## 2 Nuclear-matter equation of state

Let us first reconsider the equation of state of isospin-symmetric nuclear matter as it follows from the calculation of ref. [3]. Even though all contributions to the energy per particle  $\bar{E}(k_f)$  have been given explicitly in ref. [3] we prefer to write down again the extra terms generated by the NN-contact interactions proportional to  $g_{0,1} + g_A^2/4$  (using a more compact notation). The first diagram in fig. 1 gives rise to a contribution to the energy per particle of the form

$$\bar{E}(k_f) = -\frac{(\gamma + 1)g_A^2 k_f^3}{(4\pi f_\pi)^2}, \quad (1)$$

where we have introduced (for notational convenience) the coefficient  $\gamma$  by the relation  $(\gamma + 1)g_A^2/2 = g_0 + g_1 + g_A^2/2$ . In

the second and third diagrams in fig. 1 the contact interaction proportional to  $\gamma + 1$  is iterated with  $1\pi$ -exchange or with itself (dropping the  $\gamma^2$ -contribution). Putting a medium insertion<sup>1</sup> at each of the two nucleon propagators with equal orientation one gets

$$\bar{E}(k_f) = \frac{3(\gamma + 1)g_A^4 M m_\pi^4}{5(8\pi)^3 f_\pi^4} \left[ 11u - \frac{1}{2u} - (10 + 8u^2) \times \arctan 2u + \left( \frac{1}{8u^3} + \frac{5}{2u} \right) \ln(1 + 4u^2) \right], \quad (2)$$

with the abbreviation  $u = k_f/m_\pi$ . One observes that eq. (2) receives no contribution from the third diagram in fig. 1 since  $\int_0^\infty dl$  is set to zero in dimensional regularization. The second and third diagrams in fig. 1 with three medium insertions give rise to the following contribution to the energy per particle:

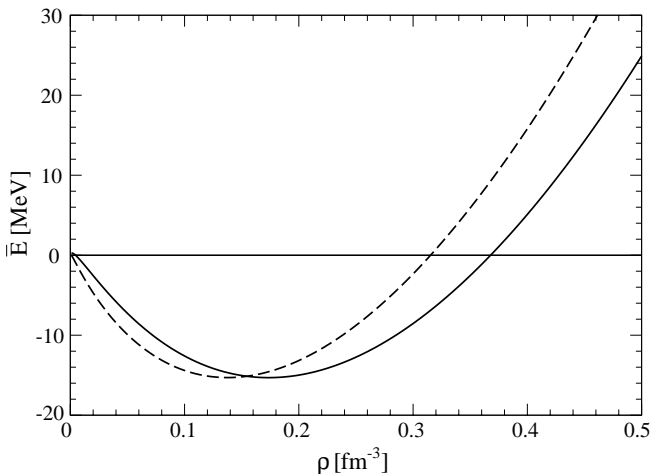
$$\bar{E}(k_f) = \frac{9g_A^4 M m_\pi^4}{(4\pi f_\pi)^4 u^3} \int_0^u dx x^2 \int_{-1}^1 dy \left[ 2uxy + (u^2 - x^2 y^2) H \right] \times \left[ \frac{\gamma + 1}{2} \ln(1 + s^2) - \frac{s^2}{4} \right], \quad (3)$$

with the auxiliary functions  $H = \ln(u + xy) - \ln(u - xy)$  and  $s = xy + \sqrt{u^2 - x^2 + x^2 y^2}$ . In the chiral limit  $m_\pi = 0$  only the contribution coming from the last term,  $-s^2/4$ , in the second square bracket survives. The corresponding double integral  $\int_0^u dx x^2 \int_{-1}^1 dy \dots$  has the value  $2u^7(\ln 4 - 11)/105$ . The expansion of the energy per particle up to order  $\mathcal{O}(k_f^4)$  is completed by adding to the terms eqs. (1)-(3) the contributions from the (relativistically improved) kinetic energy, from  $1\pi$ -exchange and from iterated  $1\pi$ -exchange written down in eqs. (5)-(11) of ref. [4]. In case of the  $1\pi$ -exchange contribution (eq. (6) in ref. [4]) we neglect, of course, the small relativistic  $1/M^2$ -correction of order  $\mathcal{O}(k_f^5)$ .

Now, we have to fix parameters. The pion decay constant  $f_\pi = 92.4$  MeV and the nucleon mass  $M = 939$  MeV are well known. As in ref. [4] we choose the value  $g_A = 1.3$ . This corresponds via the Goldberger-Treiman relation to a  $\pi$ NN-coupling constant of  $g_{\pi N} = g_A M / f_\pi = 13.2$  which agrees with present empirical determinations of  $g_{\pi N}$  from  $\pi$ N-dispersion relation analyses [8]. We set  $m_\pi = 135$  MeV (the neutral pion mass) since this is closest to the expected value of the pion mass in the absence of isospin-breaking and electromagnetic effects.

The dashed line in fig. 2 shows the equation of state of isospin-symmetric nuclear matter in the approach of ref. [3] using the above-mentioned input parameters. The coefficient  $\gamma = 4.086$  has been adjusted such that the minimum of the saturation curve lies at  $\bar{E}(k_{f0}) = -15.3$  MeV [9]. The predicted equilibrium density  $\rho_0 = 0.138$  fm<sup>-3</sup> (corresponding to a Fermi

<sup>1</sup> This is a technical notation for the difference between the in-medium and vacuum nucleon propagator. For further details, see sect. 2 in ref. [4].



**Fig. 2.** The energy per particle of isospin-symmetric nuclear matter  $\bar{E}(k_f)$  versus the nucleon density  $\rho = 2k_f^3/3\pi^2$ . The dashed line corresponds to the approach of ref. [3]. The full line results if the NN-contact interaction is treated in mean-field approximation. In each case the coefficient  $\gamma$  is adjusted such that the saturation minimum lies at  $\bar{E}(k_{f0}) = -15.3$  MeV.

momentum of  $k_{f0} = 250.1$  MeV) is somewhat too low. The same holds for the nuclear-matter compressibility  $K = k_{f0}^2 \bar{E}''(k_{f0}) = 202$  MeV. Of course, if we use the input parameters of ref. [3] ( $f_\pi = 93$  MeV,  $g_A = 1.26$ ,  $m_\pi = 140$  MeV and  $g_0 + g_1 = 3.23$  corresponding to  $\gamma = 4.07$ ) we exactly reproduce the numerical results of that work. We emphasize that the different treatment of the two components of the NN-contact interaction is essential in order to get (realistic) nuclear binding and saturation in the framework of ref. [3]. If both components were treated on equal footing in first order (technically this is realized by deleting the contribution coming from the third diagram in fig. 1) the energy per particle  $\bar{E}(k_f)$  would not even develop a minimum.

### 3 Real single-particle potential

Next, we turn to the real part of the single-particle potential  $U(p, k_f)$  below the Fermi surface ( $p \leq k_f$ ) in the framework of ref. [3]. As outlined in ref. [5], the contributions to  $U(p, k_f)$  can be classified as two-body and three-body potentials. From the first diagram in fig. 1 one gets a contribution to the two-body potential of the form

$$U_2(p, k_f) = -\frac{2(\gamma+1)g_A^2 k_f^3}{(4\pi f_\pi)^2}, \quad (4)$$

which is just twice its contribution to the energy per particle (see eq. (1)). From the second diagram in fig. 1 one derives a contribution to the two-body potential of the

form

$$U_2(p, k_f) = \frac{(\gamma+1)g_A^4 M m_\pi^4}{(4\pi)^3 f_\pi^4} \times \left\{ u + \frac{1}{4x}(x^3 - 3x - 3u^2x - 2u^3) \arctan(u+x) + \frac{1}{4x}(x^3 - 3x - 3u^2x + 2u^3) \arctan(u-x) + \frac{1}{8x}(1 + 3u^2 - 3x^2) \ln \frac{1 + (u+x)^2}{1 + (u-x)^2} \right\}, \quad (5)$$

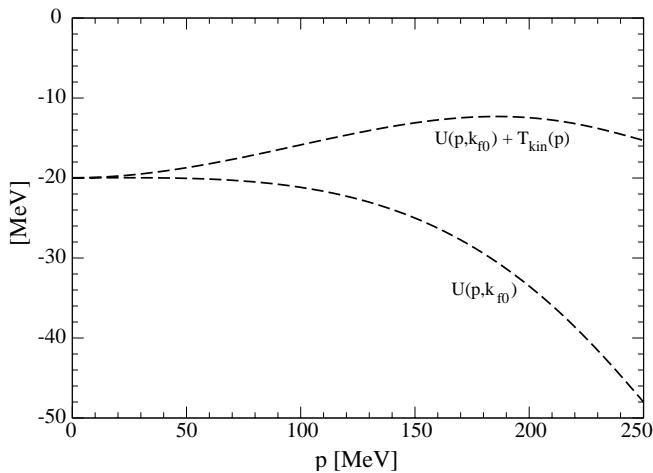
with the abbreviation  $x = p/m_\pi$ . The second and third diagrams in fig. 1 give each rise to three different contributions to the three-body potential. Altogether, they read

$$U_3(p, k_f) = \frac{3g_A^4 M m_\pi^4}{(4\pi f_\pi)^4} \times \int_{-1}^1 dy \left\{ [2uxy + (u^2 - x^2y^2)H] \times \left[ \frac{\gamma+1}{2} \ln(1+s^2) - \frac{s^2}{4} \right] + \int_{-xy}^{s-xy} d\xi \left[ 2u\xi + (u^2 - \xi^2) \ln \frac{u+\xi}{u-\xi} \right] \times \frac{(2\gamma+1)(xy+\xi) - (xy+\xi)^3}{2[1+(xy+\xi)^2]} + \int_0^u d\xi \frac{\xi^2}{x} \left[ (\gamma+1) \ln(1+\sigma^2) - \frac{\sigma^2}{2} \right] \ln \frac{|x+\xi y|}{|x-\xi y|} \right\}, \quad (6)$$

with the auxiliary function  $\sigma = \xi y + \sqrt{u^2 - \xi^2 + \xi^2 y^2}$ . The real single-particle potential  $U(p, k_f)$  is completed by adding to the terms eqs. (4)-(6) the contributions from  $1\pi$ -exchange and iterated  $1\pi$ -exchange written down in eqs. (8)-(13) of ref. [5]. Again, the (higher-order) relativistic  $1/M^2$ -correction to  $1\pi$ -exchange (see eq. (8) in ref. [5]) is neglected for reasons of consistency.

The lower curve in fig. 3 shows the momentum dependence of the real single-particle potential  $U(p, k_{f0})$  at saturation density  $k_{f0} = 250.1$  MeV as it arises in the framework of Lutz *et al.* [3]. The predicted potential depth for a nucleon at the bottom of the Fermi sea is only  $U(0, k_{f0}) = -20.0$  MeV. In magnitude this is much smaller than the typical depth  $U_0 \simeq -53$  MeV of the empirical optical model potential [10] or the nuclear shell model potential [11]. The upper curve in fig. 3 includes the (relativistically improved) single-nucleon kinetic energy  $T_{\text{kin}}(p) = p^2/2M - p^4/8M^3$ . As required by the Hugenholtz-van-Hove theorem [12] this curve ends at the Fermi surface  $p = k_{f0}$  with the value  $\bar{E}(k_{f0}) = -15.3$  MeV. A further important check is provided by the sum rule for the two- and three-body potentials  $U_{2,3}(p, k_f)$  written down in eq. (5) of ref. [5]. It holds with very high numerical accuracy in the present calculation.

The momentum dependence of the two (dashed) curves in fig. 3 is completely unrealistic. Most seriously, the total single-particle energy  $T_{\text{kin}}(p) + U(p, k_{f0})$  (upper curve) does not rise monotonically with the nucleon momentum  $p$ , but instead it starts to bend downward above

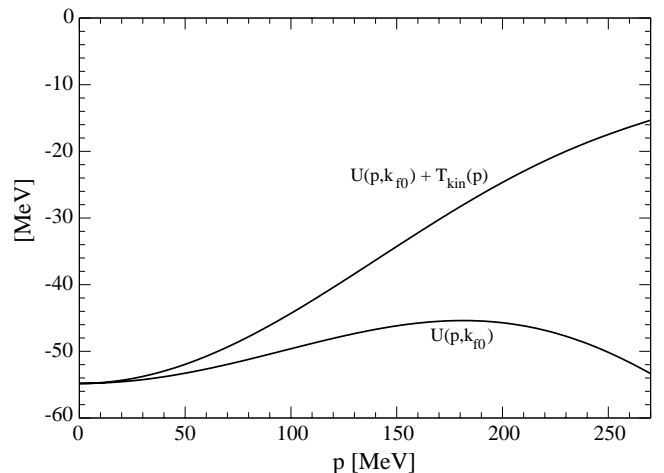


**Fig. 3.** The lower curve shows the real part of the single-particle potential  $U(p, k_{f0})$  at saturation density  $k_{f0} = 250.1$  MeV in the approach of Lutz *et al.* [3]. The upper curve includes in addition the relativistically improved kinetic energy  $T_{\text{kin}}(p) = p^2/2M - p^4/8M^3$ .

$p \simeq 190$  MeV. This implies a negative effective nucleon mass at the Fermi surface,  $M^*(k_{f0}) \simeq -3.5 M$ , and a negative density of states with dramatic consequences for the finite-temperature behavior of nuclear matter. Because of such unrealistic features of the underlying single-particle potential, the scheme of Lutz *et al.* [3] has to be modified substantially.

The overly strong momentum dependence of  $U(p, k_{f0})$  comes from the second and third diagrams in fig. 1 in which the NN-contact interaction proportional to the large coefficient  $\gamma+1$  is further iterated. In order to demonstrate this feature we drop these (iterated) three-loop diagrams and keep the NN-contact interaction (of unspecified dynamical origin) at the mean-field level. The resulting equation of state obtained by leaving out the contributions eqs. (2), (3) and adjusting  $\gamma = 6.198$  is shown by the full line in fig. 2. The predicted saturation density is now  $\rho_0 = 0.174 \text{ fm}^{-3}$  (corresponding to a Fermi momentum of  $k_{f0} = 270.3$  MeV) and the nuclear-matter compressibility has the value  $K = 253$  MeV. We emphasize that the scheme of ref. [3] modified by a mean-field treatment of the NN-contact interaction becomes fully equivalent to the truncation at fourth order in small momenta of our previous work [4, 5] after the identification of parameters,  $\gamma + 1 = 10g_A^2 \Lambda M / (4\pi f_\pi)^2$ , with  $\Lambda \simeq 0.61$  GeV denoting the cut-off scale. We remind that in ref. [4] the necessary short-ranged attractive contribution was not introduced explicitly, but generated at second order by iterated  $1\pi$ -exchange employing (the more physical) cut-off regularization.

The lower full curve in fig. 4 shows the momentum dependence of the real single-particle potential at saturation density  $k_{f0} = 270.3$  MeV which results in a mean-field approximation of the NN-contact interaction (by leaving out the contributions eqs. (5), (6)). The predicted potential depth  $U(0, k_{f0}) = -54.8$  MeV is in good agreement



**Fig. 4.** The lower curve shows the real part of the single-particle potential  $U(p, k_{f0})$  at saturation density  $k_{f0} = 270.3$  MeV in a mean-field treatment of the NN-contact interaction.

with that of optical model [10] or nuclear shell model potentials [11]. Most importantly, the total single-particle energy  $T_{\text{kin}}(p) + U(p, k_{f0})$  (upper curve) grows now monotonically with the nucleon momentum  $p$ , as it should. The up- and downward bending of the lower full curve in fig. 4 is however still too strong. The negative slope of  $U(p, k_{f0})$  at the Fermi surface  $p = k_{f0}$  leads to a too large effective nucleon mass  $M^*(k_{f0}) \simeq 2.9 M$  which reflects itself in a too high critical temperature  $T_c \simeq 25$  MeV of the liquid-gas phase transition [7].

Another possibility (closer in spirit to the original proposal of Lutz *et al.* [3, 6]) is to keep the (large) components of the NN-contact interaction proportional to  $g_{0,1}$  at the mean-field level and to iterate only the (small) components proportional to  $g_A^2/4$ . Such a “partial mean-field approximation” is realized by simply setting  $\gamma = 0$  in eqs. (2), (3), (5), (6). In fact this approach is fully equivalent to calculating the  $1\pi$ -exchange Fock diagram and the iterated  $1\pi$ -exchange Hartree and Fock diagrams with the “regularized”  $1\pi$ -exchange NN  $T$ -matrix of the form

$$\mathcal{T}_{\text{NN}}^{(1\pi)} = \frac{g_A^2}{4f_\pi^2} \left[ \frac{\vec{\sigma}_1 \cdot \vec{q} \vec{\sigma}_2 \cdot \vec{q}}{m_\pi^2 + \vec{q}^2} - \frac{\vec{\sigma}_1 \cdot \vec{\sigma}_2}{3} \right] \vec{\tau}_1 \cdot \vec{\tau}_2. \quad (7)$$

Here,  $\vec{\sigma}_{1,2}$  and  $\vec{\tau}_{1,2}$  denote, as usual, the spin and isospin operators of the two nucleons and  $\vec{q}$  is the momentum transfer between both nucleons carried by the exchanged pion. When using the  $T$ -matrix, eq. (7) the zero-range pieces generated by  $1\pi$ -exchange are removed order by order as proposed in ref. [6].

After adjusting the coefficient  $\gamma = 5.363$  to the maximum binding energy  $-\bar{E}(k_{f0}) = 15.3$  MeV one finds in this “partial mean-field approximation” an equilibrium density of  $\rho_0 = 0.190 \text{ fm}^{-3}$  (corresponding to a Fermi momentum of  $k_{f0} = 278.3$  MeV) which lies somewhat too high. At the same time the nuclear compressibility comes

out as  $K = 248$  MeV and the potential depth increases (in magnitude) to  $U(0, k_{f0}) = -63.5$  MeV. A very welcome side effect of this “partial mean-field approximation” is a weaker  $p$ -dependence of the real single-particle potential  $U(p, k_{f0})$  implying a more realistic value of the effective nucleon mass at the Fermi surface of  $M^*(k_{f0}) \simeq 1.5 M$ . Furthermore, if one employs cut-off regularization (instead of dimensional regularization) the adjustable parameter  $\gamma = 5.363$  can be interpreted in terms of a momentum cut-off  $\Lambda \simeq 0.57$  GeV by making use of the relation  $\gamma = 2g_A^2 \Lambda M / (2\pi f_\pi)^2$ .

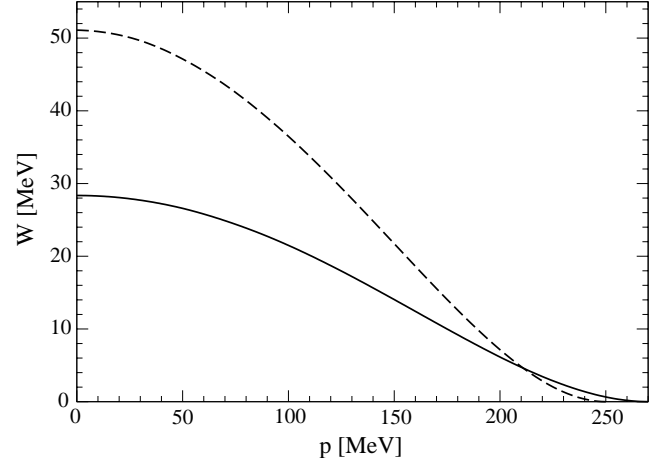
#### 4 Imaginary single-particle potential

In this section, we discuss the imaginary part of the single-particle potential  $W(p, k_f)$  for  $p \leq k_f$  as it arises in the scheme of Lutz *et al.* [3]. This quantity determines the half-width of nucleon hole states in the Fermi sea. As outlined in ref. [5] the contributions to  $W(p, k_f)$  can be classified as two-body, three-body and four-body terms. From the second and third diagrams in fig. 1 one derives a two-body term of the form

$$W_2(p, k_f) = \frac{g_A^4 M m_\pi^4}{(8\pi)^3 f_\pi^4} \times \left\{ u^2 x^2 + \frac{3u^4}{2} - \frac{x^4}{10} + (\gamma + 1) \left[ 4 + 14u^2 - \frac{22x^2}{3} + \frac{2}{x}(3x^2 - 3u^2 - 1) \left[ \arctan(u+x) - \arctan(u-x) \right] + \frac{1}{x}(x^3 - 3x - 3u^2x - 2u^3) \ln[1 + (u+x)^2] + \frac{1}{x}(x^3 - 3x - 3u^2x + 2u^3) \ln[1 + (u-x)^2] \right] \right\}. \quad (8)$$

The associated three-body term reads

$$W_3(p, k_f) = \frac{g_A^4 M m_\pi^4}{(8\pi)^3 f_\pi^4} \left\{ 2x^4 - 6u^4 + (\gamma + 1) \times \left[ \frac{41x^2}{3} - 31u^2 - 5 - (u^2 - x^2)^2 - 3 \ln(1 + 4x^2) + \left( 6x - \frac{3}{2x} \right) \arctan 2x + \left[ \arctan(u+x) - \arctan(u-x) \right] \times \frac{1}{2x} \left[ (u^2 - x^2)^3 + (12u^2 + 27)(u^2 - x^2) + 8 \right] + \left( 6 + 9u^2 + \frac{2u^3}{x} - 3x^2 \right) \ln[1 + (u+x)^2] + \left( 6 + 9u^2 - \frac{2u^3}{x} - 3x^2 \right) \ln[1 + (u-x)^2] \right] \right\}, \quad (9)$$



**Fig. 5.** The imaginary part of the single-particle potential  $W(p, k_{f0})$  at saturation density *versus* the nucleon momentum  $p$ . The dashed line corresponds to the approach of Lutz *et al.* [3] and the full line shows the result obtained in a mean-field treatment of the NN-contact interaction.

and the four-body term is given by the expression

$$W_4(p, k_f) = \frac{g_A^4 M m_\pi^4}{(8\pi)^3 f_\pi^4} \left\{ 3u^4 + 2u^2x^2 - \frac{17x^4}{5} + (\gamma + 1) \left[ 1 + 20u^2 - \frac{28x^2}{3} + (u^2 - x^2)^2 + 6 \ln(1 + 4x^2) + \left( \frac{3}{x} - 12x \right) \arctan 2x + \left[ \arctan(u+x) - \arctan(u-x) \right] \times \frac{1}{2x} \left[ 3x^4 + 6u^2x^2 - 9u^4 - (u^2 - x^2)^3 + 27x^2 - 15u^2 - 7 \right] + 2(x^2 - 3 - 3u^2) \times \left( \ln[1 + (u+x)^2] + \ln[1 + (u-x)^2] \right) \right] \right\}. \quad (10)$$

The additional contributions from the iterated  $1\pi$ -exchange Hartree and Fock diagram are collected in eqs. (20)-(25) of ref. [5]. The total imaginary single-particle potential evaluated at zero nucleon momentum ( $p = 0$ ) can even be written as a closed form expression

$$W(0, k_f) = \frac{3\pi g_A^4 M m_\pi^4}{(4\pi f_\pi)^4} \left\{ \frac{u^4}{2} + (\gamma - 2)u^2 - \frac{2u^2}{1+u^2} + \frac{\pi^2}{12} + \text{Li}_2(-1-u^2) + \left[ 4 - \gamma + \ln(2+u^2) - \frac{1}{2} \ln(1+u^2) \right] \ln(1+u^2) \right\}, \quad (11)$$

where  $\text{Li}_2(-a^{-1}) = \int_0^1 d\zeta (\zeta + a)^{-1} \ln \zeta$  denotes the conventional dilogarithmic function.

The dashed line in fig. 5 shows the momentum dependence of the imaginary single-particle potential  $W(p, k_{f0})$

at saturation density  $k_{f0} = 250.1$  MeV as it arises in the approach of ref. [3]. The predicted value  $W(0, k_{f0}) = 51.1$  MeV lies outside the range 20–40 MeV obtained in calculations based on (semi)-realistic NN-forces [13, 14]. The full line in fig. 5 corresponds to a mean-field approximation of the NN-contact interaction. Up to a slight change in the equilibrium Fermi momentum  $k_{f0} = 270.3$  MeV the full curve in fig. 5 agrees with the one shown in fig. 4 of ref. [5]. The considerably reduced value  $W(0, k_{f0}) = 28.4$  MeV indicates the large contribution of the iterated diagrams in fig. 1 to the imaginary single-particle potential  $W(p, k_f)$ . Note that both curves in fig. 5 vanish quadratically near the Fermi surface as required by Luttinger’s theorem [15]. As further check on our calculation we verified the zero sum rule:  $\int_0^{k_f} dp p^2 [6W_2(p, k_f) + 4W_3(p, k_f) + 3W_4(p, k_f)] = 0$ , for the two-, three- and four-body components  $W_{2,3,4}(p, k_f)$  written in eqs. (8)-(10). In the “partial mean-field approximation” (realized by setting  $\gamma = 0$  in eqs. (8)-(11)) the imaginary single-particle potential gets also substantially reduced as indicated by the value at zero nucleon momentum,  $W(0, k_{f0} = 278.3$  MeV) = 24.0 MeV.

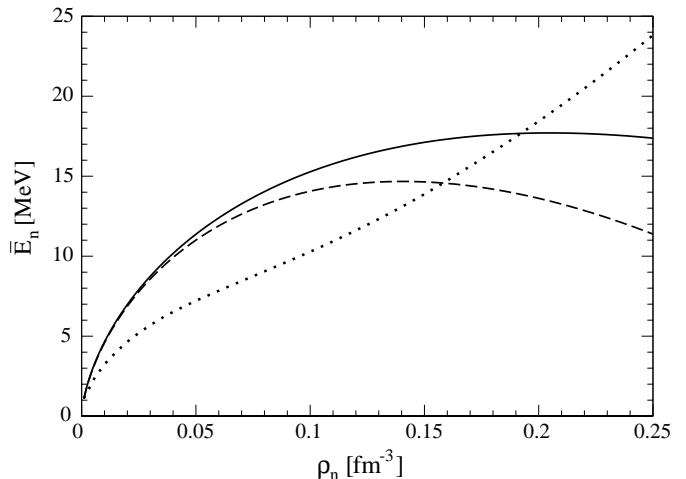
## 5 Neutron matter

In this section we discuss the equation of state of pure neutron matter. In the scheme of Lutz *et al.* [3] the energy per particle of pure neutron matter  $\bar{E}_n(k_n)$  depends exclusively on the coefficient  $g_1$  parameterizing the short-range NN-interaction in the channel with total isospin  $I = 1$ . There is no need to write down explicitly the contributions of the diagrams in fig. 1 to  $\bar{E}_n(k_n)$ . These expressions are easily obtained from eqs. (1)-(3) by replacing  $k_f$  by the neutron Fermi momentum  $k_n$ , by replacing the coefficient  $\gamma$  by a new one  $\gamma_n$ , and by multiplying the formulas with a relative isospin factor  $1/3$ . The relation  $(\gamma_n + 1)g_A^2/4 = g_1 + g_A^2/4$  defines this new coefficient  $\gamma_n$ . The additional contributions to  $\bar{E}_n(k_n)$  from the kinetic energy,  $1\pi$ -exchange and iterated  $1\pi$ -exchange are written down in eqs. (32)-(37) of ref. [4] (neglecting again the relativistic  $1/M^2$ -correction to  $1\pi$ -exchange).

The dashed line in fig. 6 shows the energy per particle of pure neutron matter  $\bar{E}_n(k_n)$  versus the neutron density  $\rho_n = k_n^3/3\pi^2$  as it arises in the approach of ref. [3]. The coefficient  $\gamma_n = 0.055$  has been adjusted to the empirical value of the asymmetry energy  $A(k_{f0} = 250.1$  MeV) = 33.2 MeV (see next section). The downward bending of the dashed curve in fig. 6 above  $\rho_n > 0.15$  fm $^{-3}$  is even stronger than in our previous work [4] (see fig. 8 therein). This property can be understood by taking the chiral limit ( $m_\pi \rightarrow 0$ ) of the calculated neutron matter equation of state and considering the coefficient  $\beta_n$  in front of the term  $k_n^4/M^3$ . In the approach of Lutz *et al.* [3] one has

$$\beta_n = -\frac{1}{70} \left( \frac{g_{\pi N}}{4\pi} \right)^4 (4\pi^2 + 17 + 16 \ln 2) - \frac{3}{56} = -1.23, \quad (12)$$

which is 2.2 times the negative value of  $\beta_n$  found in ref. [4]. The full line in fig. 6 shows the equation of state of pure



**Fig. 6.** The energy per particle of pure neutron matter  $\bar{E}_n(k_n)$  versus the neutron density  $\rho_n = k_n^3/3\pi^2$ . The dashed line corresponds to the approach of Lutz *et al.* [3] and the full line shows the result obtained in mean-field approximation of the  $nn$ -contact interaction. The dotted line stems from the many-body calculation of the Urbana group [16].

neutron matter obtained in the mean-field approximation of the  $nn$ -contact interaction proportional to  $\gamma_n + 1$  after adjusting  $\gamma_n = 0.788$  to the empirical value of the asymmetry energy  $A(k_{f0} = 270.3$  MeV) = 33.2 MeV. The downward bending of the full curve in fig. 6 is weaker and it sets in at somewhat higher densities  $\rho_n > 0.2$  fm $^{-3}$ . The dotted line in fig. 6 stems from the many-body calculation of the Urbana group [16]. This curve should be considered as a representative of the host of existing realistic neutron matter calculations which scatter around it. We also note that the result for  $\bar{E}_n(k_n)$  obtained in the “partial mean-field approximation” (after adjusting  $\gamma_n = -0.131$ ) is very similar to the full line in fig. 6 for neutron densities  $\rho_n < 0.25$  fm $^{-3}$ . The systematic deviations observed in fig. 6 indicate that the neutron matter equation of state of ref. [4] cannot be improved by treating the short-range NN-dynamics as proposed in ref. [3].

## 6 Asymmetry energy

Finally, we turn to the density-dependent asymmetry energy  $A(k_f)$  in the approach of ref. [3]. The asymmetry energy is generally defined by the expansion of the energy per particle of isospin-asymmetric nuclear matter (described by different proton and neutron Fermi momenta  $k_{p,n} = k_f(1 \mp \delta)^{1/3}$ ) around the symmetry line:  $\bar{E}_{\text{as}}(k_p, k_n) = \bar{E}(k_f) + \delta^2 A(k_f) + \mathcal{O}(\delta^4)$ . Evaluation of the first diagram in fig. 1 leads to the following contribution to the asymmetry energy:

$$A(k_f) = \frac{g_A^2 k_f^3}{3(4\pi f_\pi)^2} (3\gamma - 2\gamma_n + 1), \quad (13)$$

with the coefficients  $\gamma = 2(g_0 + g_1)/g_A^2$  and  $\gamma_n = 4g_1/g_A^2$  in the notation of ref. [3]. Putting a medium insertion at

each of two nucleon propagators with equal orientation one gets from the second and third diagrams in fig. 1

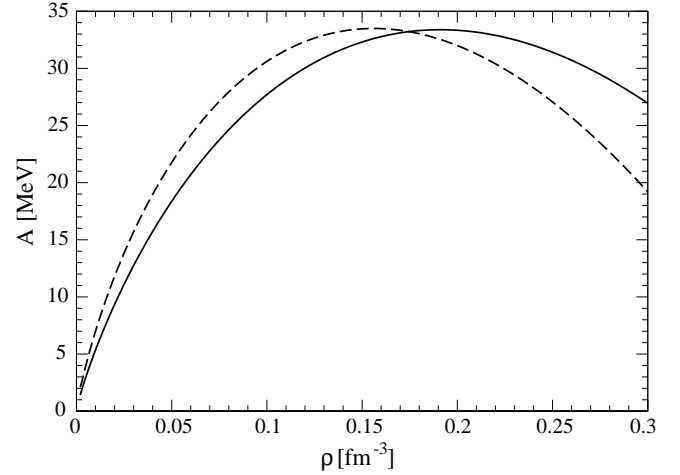
$$A(k_f) = \frac{g_A^4 M m_\pi^4}{3(8\pi)^3 f_\pi^4} \left\{ 2(\gamma+1)u + 8(2\gamma-\gamma_n+1)u^2 \arctan 2u \right. \\ \left. + \left[ (2\gamma_n-6\gamma-4)u - \frac{\gamma+1}{2u} \right] \ln(1+4u^2) \right\}. \quad (14)$$

The same diagrams with three medium insertions give rise to the following contribution to the asymmetry energy:

$$A(k_f) = \frac{g_A^4 M m_\pi^4}{(4\pi f_\pi)^4 u^3} \int_0^u dx x^2 \int_{-1}^1 dy \left\{ \left[ 2uxy + (u^2 - x^2 y^2) H \right] \right. \\ \times \left( 4ss' - \frac{2}{3}s'^2 - \frac{2}{3}ss'' - \frac{7}{2}s^2 \right) \\ \left. + (\gamma+1) \left\{ \left[ \frac{uxy(11u^2 - 15x^2 y^2)}{3(u^2 - x^2 y^2)} \right. \right. \right. \\ \left. \left. + \frac{1}{2}(u^2 - 5x^2 y^2) H \right] \ln(1+s^2) - \frac{4u^2 s^2 H}{3(1+s^2)} \right. \right. \\ \left. \left. + \frac{2uxy + (u^2 - x^2 y^2) H}{6(1+s^2)^2} \right. \right. \\ \left. \times \left[ 8s(1+s^2)(3s+s''-5s') \right. \right. \\ \left. \left. + (1-s^2)(3s^2 - 8ss' + 8s'^2) \right] \right\} \\ \left. + 2u^2(\gamma_n+1) \left[ \frac{2uxy \ln(1+s^2)}{3(u^2 - x^2 y^2)} \right. \right. \\ \left. \left. + \left( \ln(1+s^2) + \frac{2s^2}{3(1+s^2)} \right) H \right] \right\}, \quad (15)$$

with  $s' = u \partial s / \partial u$  and  $s'' = u^2 \partial^2 s / \partial u^2$  denoting partial derivatives. In the chiral limit  $m_\pi = 0$  only the terms in the first and second line of eq. (15) survive. The corresponding double integral  $\int_0^u dx x^2 \int_{-1}^1 dy \dots$  has the value  $4u^7(\ln 4 - 1)/15$ . The asymmetry energy is completed by adding to the terms eqs. (13)-(15) the contributions from the kinetic energy, (static)  $1\pi$ -exchange and iterated  $1\pi$ -exchange written down in eqs. (20)-(26) of ref. [4].

The dashed line in fig. 7 shows the density dependence of the asymmetry energy  $A(k_f)$  in the approach of Lutz *et al.* [3] with the coefficient  $\gamma_n = 0.055$  adjusted (at fixed  $\gamma = 4.086$ ) to the empirical value  $A(k_{f0} = 250.1 \text{ MeV}) = 33.2 \text{ MeV}$  [9]. The full line in fig. 7 corresponds to the result obtained in mean-field approximation of the NN-contact interaction by dropping the contributions eqs. (14), (15). In that case the empirical value  $A(k_{f0} = 270.3 \text{ MeV}) = 33.2 \text{ MeV}$  [9] is reproduced by tuning (at fixed  $\gamma = 6.198$ ) the coefficient  $\gamma_n$  to the value  $\gamma_n = 0.788$ . We also note that the result for the density-dependent asymmetry energy  $A(k_f)$  in the ‘‘partial mean-field approximation’’ (realized by setting  $\gamma = 5.363$  and  $\gamma_n = -0.131$  in eq. (13) and by setting  $\gamma = \gamma_n = 0$  in eqs. (14), (15)) is almost identical to the full line in fig. 7. Both curves in fig. 7 behave rather similarly. In each case the



**Fig. 7.** The asymmetry energy  $A(k_f)$  versus the nucleon density  $\rho = 2k_f^3/3\pi^2$ . The dashed line corresponds to the approach of Lutz *et al.* [3] and the full line shows the result obtained in mean-field approximation of the NN-contact interactions. The parameter  $\gamma_n$  is in each case adjusted to the (empirical) value  $A(k_{f0}) = 33.2 \text{ MeV}$  [9].

asymmetry  $A(k_f)$  reaches its maximum close to the respective saturation density  $\rho_0$  and then it starts to bend downward. Since the same (somewhat unusual) feature has also been observed in ref. [4] it seems to be generic for perturbative chiral  $\pi$ N-dynamics truncated at three-loop order.

## 7 Concluding remarks

In this work we have continued and extended the chiral approach to nuclear matter of Lutz *et al.* [3] by calculating the underlying single-particle potential. The potential for a nucleon at the bottom of the Fermi sea  $U(0, k_{f0}) = -20.0 \text{ MeV}$  is not deep enough. Most seriously, the total single-particle energy  $T_{\text{kin}}(p) + U(p, k_{f0})$  does not grow monotonically with the nucleon momentum  $p$ . The thereof implied negative effective nucleon mass at the Fermi surface  $M^*(k_{f0}) \simeq -3.5M$  and the negative density of states will ruin the behavior of nuclear matter at finite temperatures. The half-width of nucleon holes at the bottom of the Fermi sea  $W(0, k_{f0}) = 51.1 \text{ MeV}$  comes also out too large in that approach. A good nuclear-matter equation of state and better (but still not yet optimal) single-particle properties can be obtained if the NN-contact interaction (proportional to the large coefficient  $g_0 + g_1 + g_A^2/2$ ) is kept at the mean-field level and not further iterated. The so-modified scheme is fully equivalent to a calculation without any explicit short-range NN-terms when employing cut-off regularization [4]. The energy per particle of pure neutron matter  $\bar{E}_n(k_n)$  and the asymmetry energy  $A(k_f)$  depend on a second parameter  $g_1$  in the scheme of ref. [3]. Their density dependence is similar to the results of the one-parameter calculation in ref. [4]. The downward bending of  $\bar{E}_n(k_n)$  and  $A(k_f)$  above saturation density  $\rho_0$  (less pronounced if the NN-contact interaction is kept at the mean-field level) seems to be

generic for perturbative chiral  $\pi$ N-dynamics. More elaborate calculations of nuclear matter in effective (chiral) field theory with particular attention on the short-range NN-dynamics are necessary. It is mandatory that the effective short-range NN-interaction is compatible with all (semi)-empirical properties of nuclear matter simultaneously.

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